

Some NIST Capabilities Related to CLARREO

Joseph Rice, Sergey Mekhontsev, Leonard Hanssen, Eric Shirley,
Jinan Zeng, Steve Brown, Keith Lykke, Allan Smith, Carol Johnson, Robert
Saunders, Toni Litorja, Joel Fowler, Stephen Maxwell, David Allen, Howard Yoon,
Jeanne Houston, and Jerry Fraser

Optical Technology Division
National Institute of Standards and Technology
Gaithersburg, MD 20899 USA

Note. References are made to certain commercially-available products in this presentation to adequately specify the experimental procedures involved. Such identification does not imply recommendation or endorsement by the National Institute of Standards and Technology, nor does it imply that these products are the best for the purpose specified.

Outline: Some NIST Capabilities Related to CLARREO

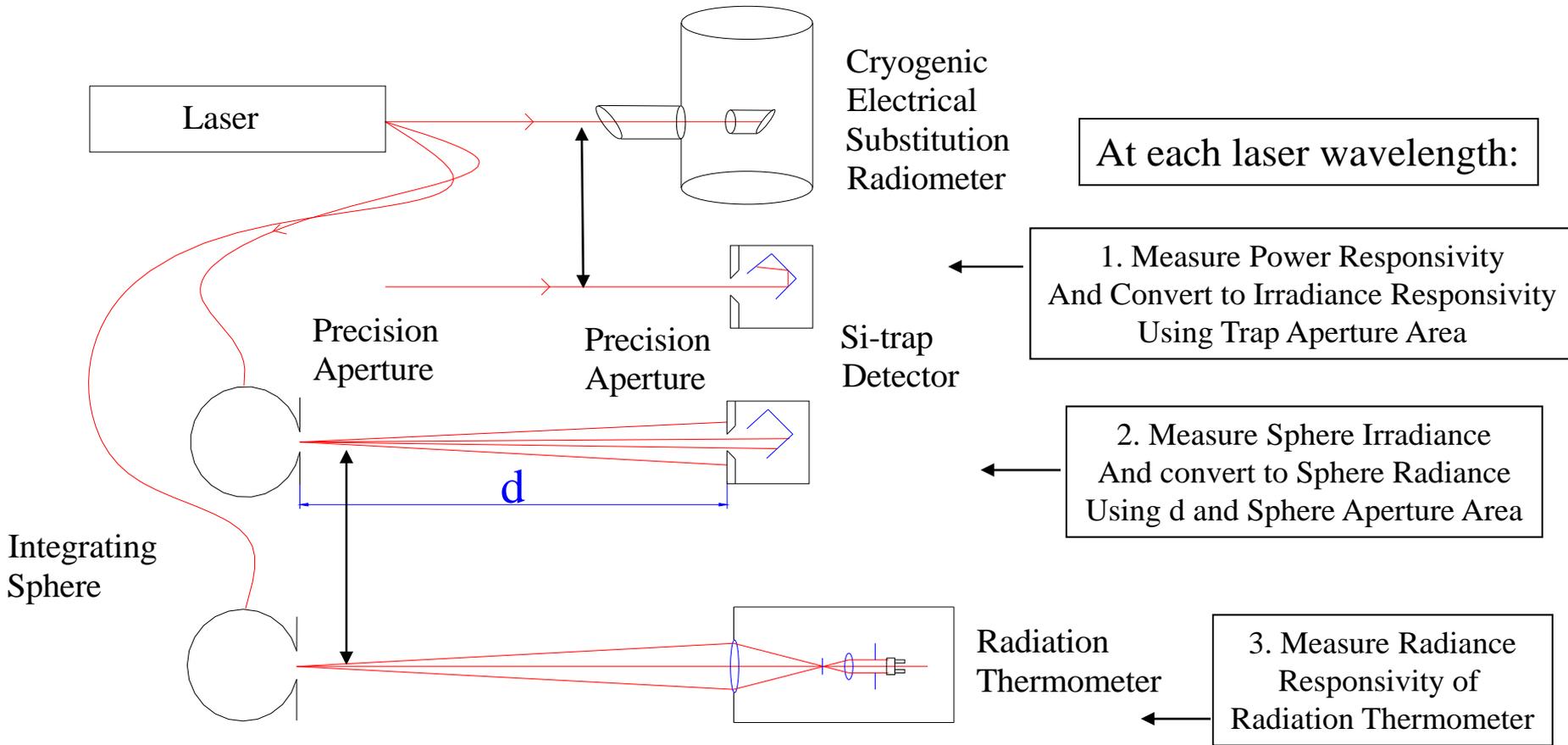
Solar-Reflected (Joe Rice, first half)

- Spectral Irradiance and Radiance Responsivity with Uniform Sources (SIRCUS)
- Hyperspectral Image Projector (HIP)
- Absolute Spectrally-Tunable Detector-Based Source

IR (Sergey Mekhontsev, 2nd half)

- Diffuse and specular reflectance, transmittance, and emittance using Fourier Transform Spectrometers
- Total-Integrated-Scatter and IR BRDF using IR lasers
- Standard and Transfer Blackbodies
- Operation in a vacuum chamber: CBS3

Establishment of the Spectral Radiance Responsivity Scale at SIRCUS

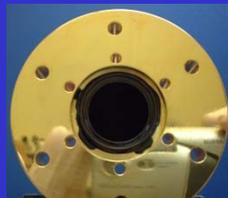
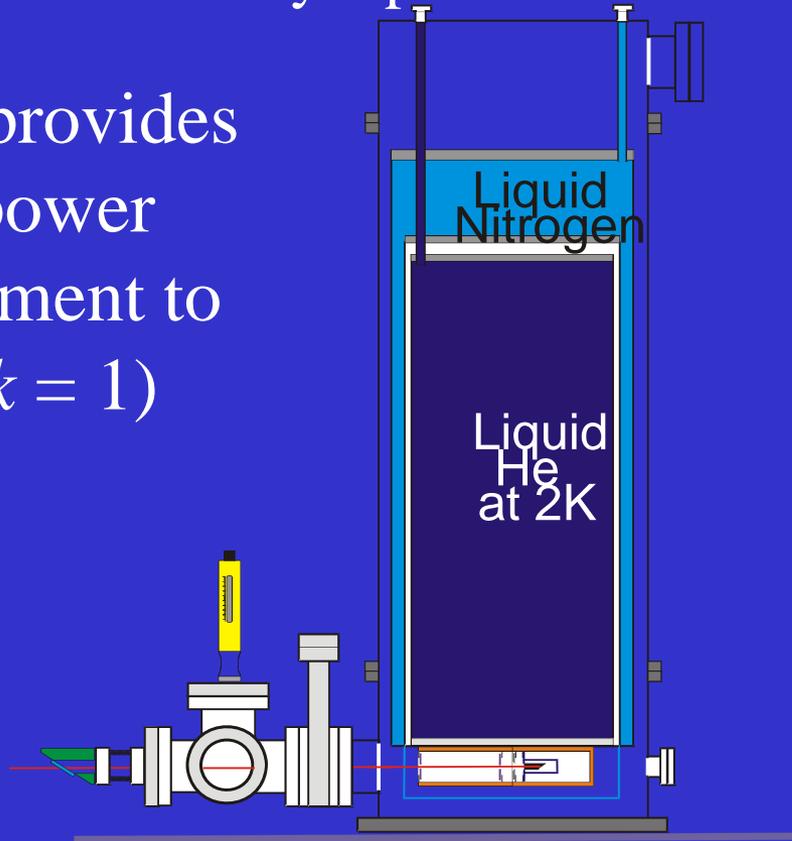


$$L(\lambda, T) = \frac{c_{1L}}{n^2 \lambda^5 \left(\exp\left(\frac{hc}{n\lambda kT}\right) - 1 \right)}$$

For realization and dissemination of temperature scales above Ag freezing point: Radiation Thermometer assigns temperature T to blackbody based on radiance L .

NIST Optical Measurements are Traceable to the Electrical Watt through the Primary Optical Watt Radiometer (POWR)

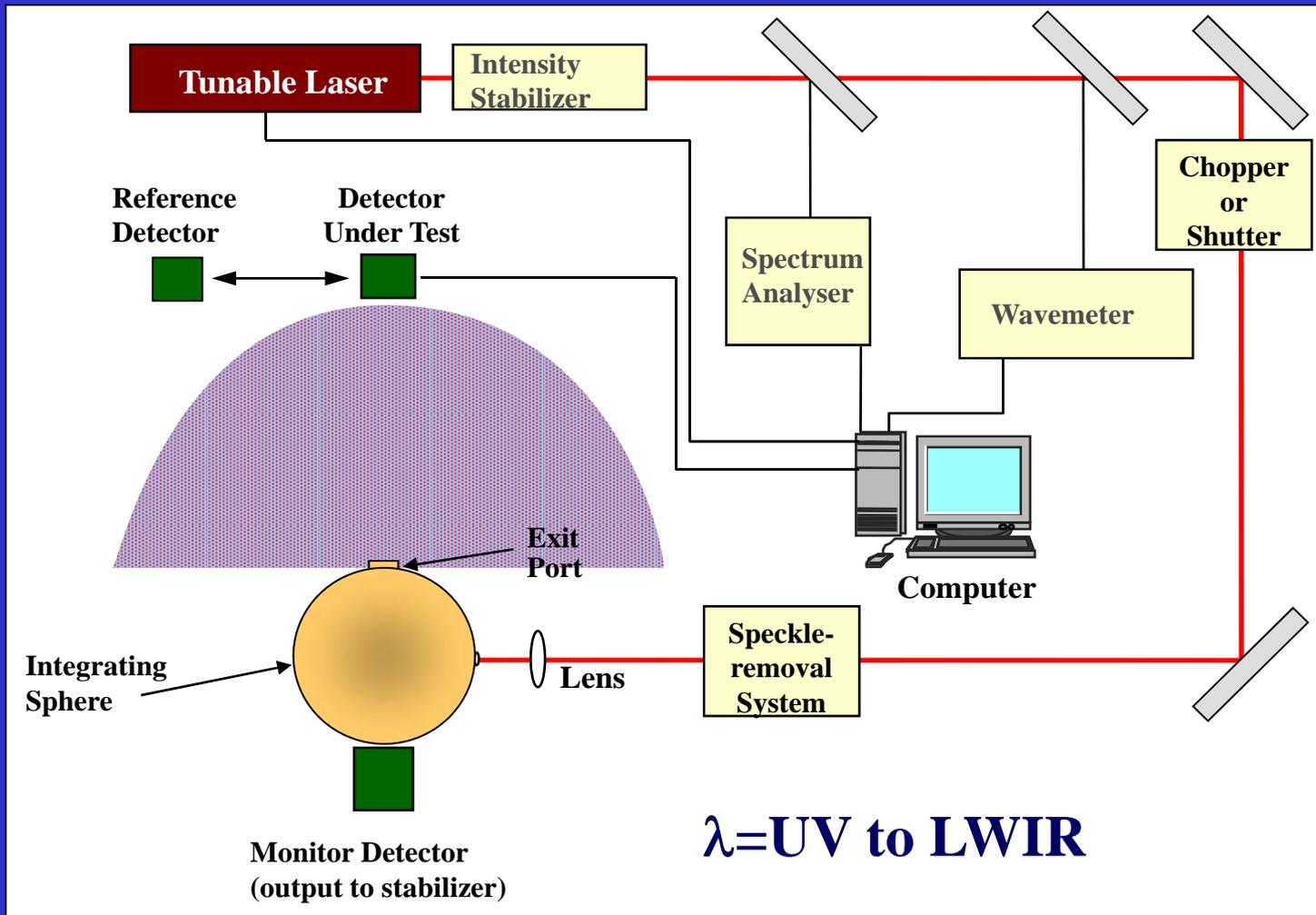
- POWR provides optical power measurement to 0.01% ($k = 1$)



*Allan Smith,
Jeanne Houston, Joe Rice*

Spectral Irradiance and Radiance Responsivity Calibrations using Uniform Sources (SIRCUS)

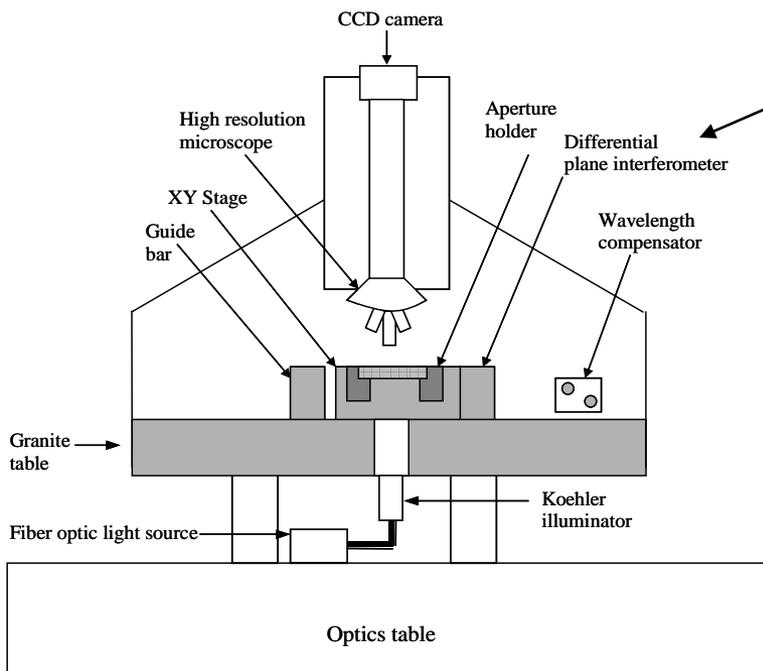
- A version of this will be implemented at NASA Goddard for use with testing the Solar Reflected CLARREO Calibration Development System



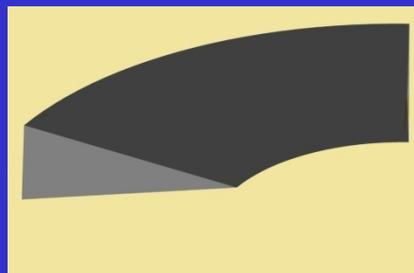
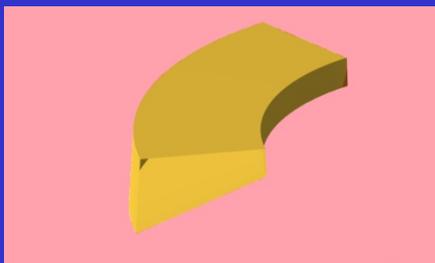
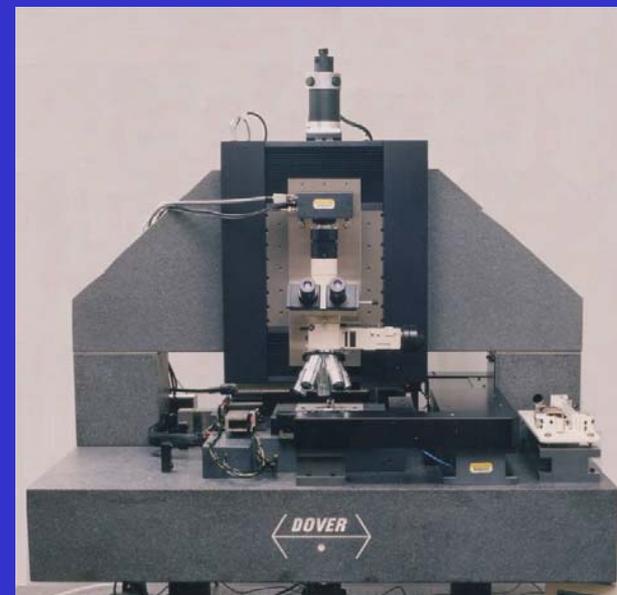
*Keith Lykke
Steve Brown
George Eppeldauer*

Aperture Area Measurements are Performed at NIST by the Absolute Aperture Area Measurement Machine

Aperture area measured to better than 0.01%

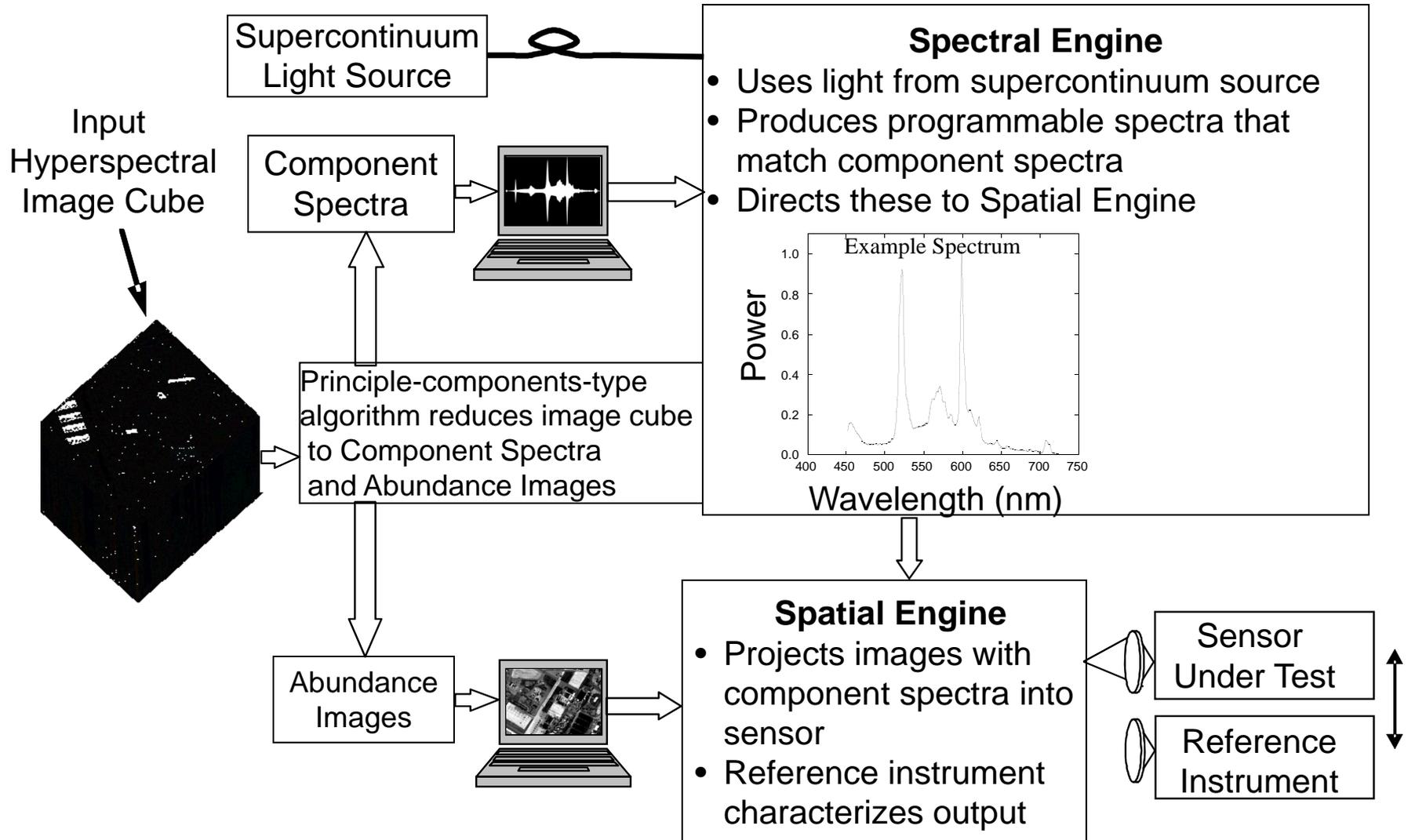


Length metrology through interferometer traceable to HeNe laser wavelength



*Toni Litorja
Joel Fowler*

Hyperspectral Image Projector (HIP) Concept

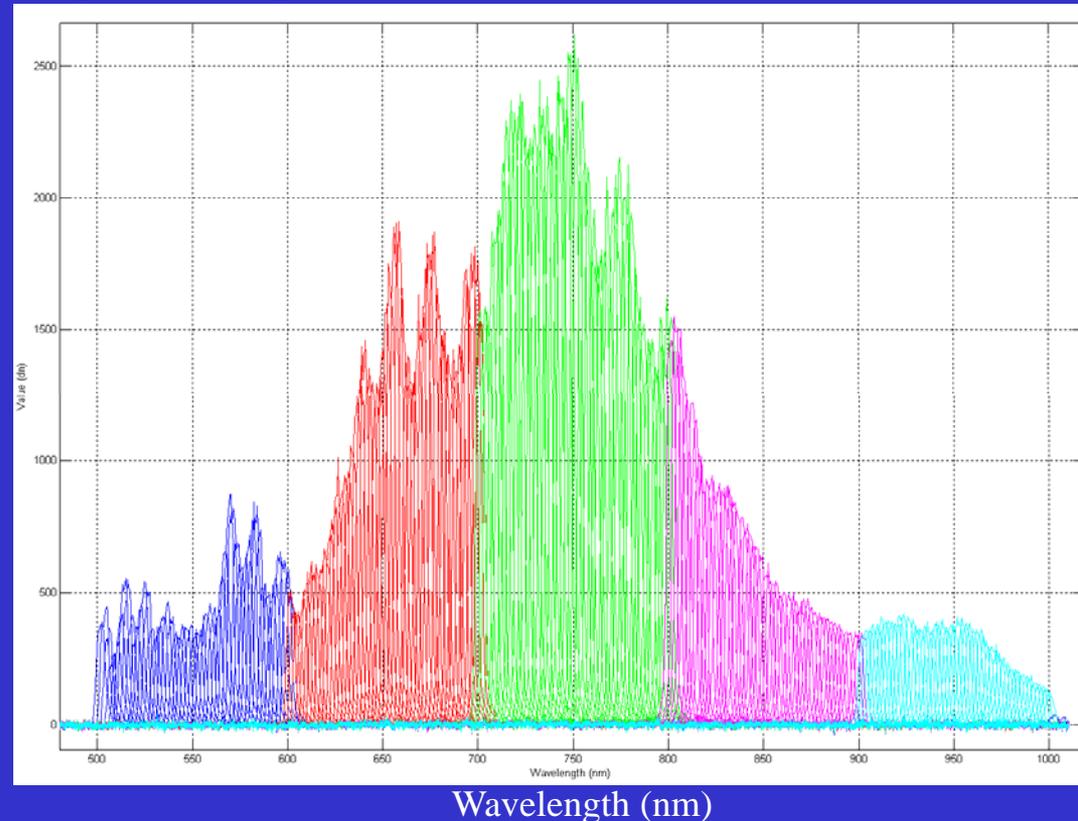


LASP IIP Hyperspectral Imager (HSI) Tests at the HIP

- “Sun” scans with solar spectrum
- “MTF” patterns test optical performance and stray light
- Simulated ground track motions of real scene
- Digital attenuation studies validate linearity and attenuations
- Spectral response calibration provides spectral and radiometric calibration

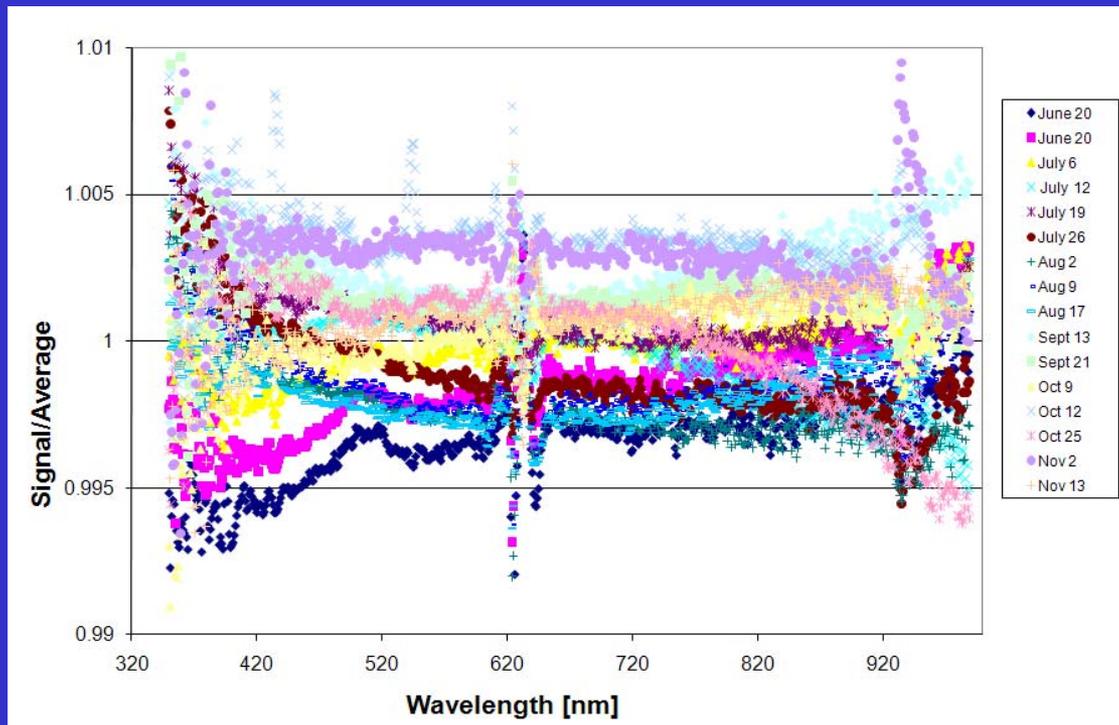


HIP Projected, LASP HSI Measured



Existing Spectral Radiance Scales Are Based on Lamps

- Laboratory solar-reflected band spectroradiometers have not demonstrated the capability that they can maintain a scale at the 0.1% level (yet). Only about 1% (1sigma) at best.
- Typical results from spectroradiometer measuring lamp-illuminated sphere:



Graphic courtesy of NIST's Remote Sensing Laboratory

- However, unfiltered radiance detectors **have demonstrated** the potential to hold a radiometric scale at the 0.1% level and better.

Absolute Spectrally-Tunable Detector-Based Source

- Absolute detector measures radiance of single-line spectra
- These data are then used to set the spectral radiance scale use to calibrate spectroradiometer

Spectrally-Tunable Source

- currently Vis, lamp-based
- could be from the HIP in future

Absolute Radiance Detector

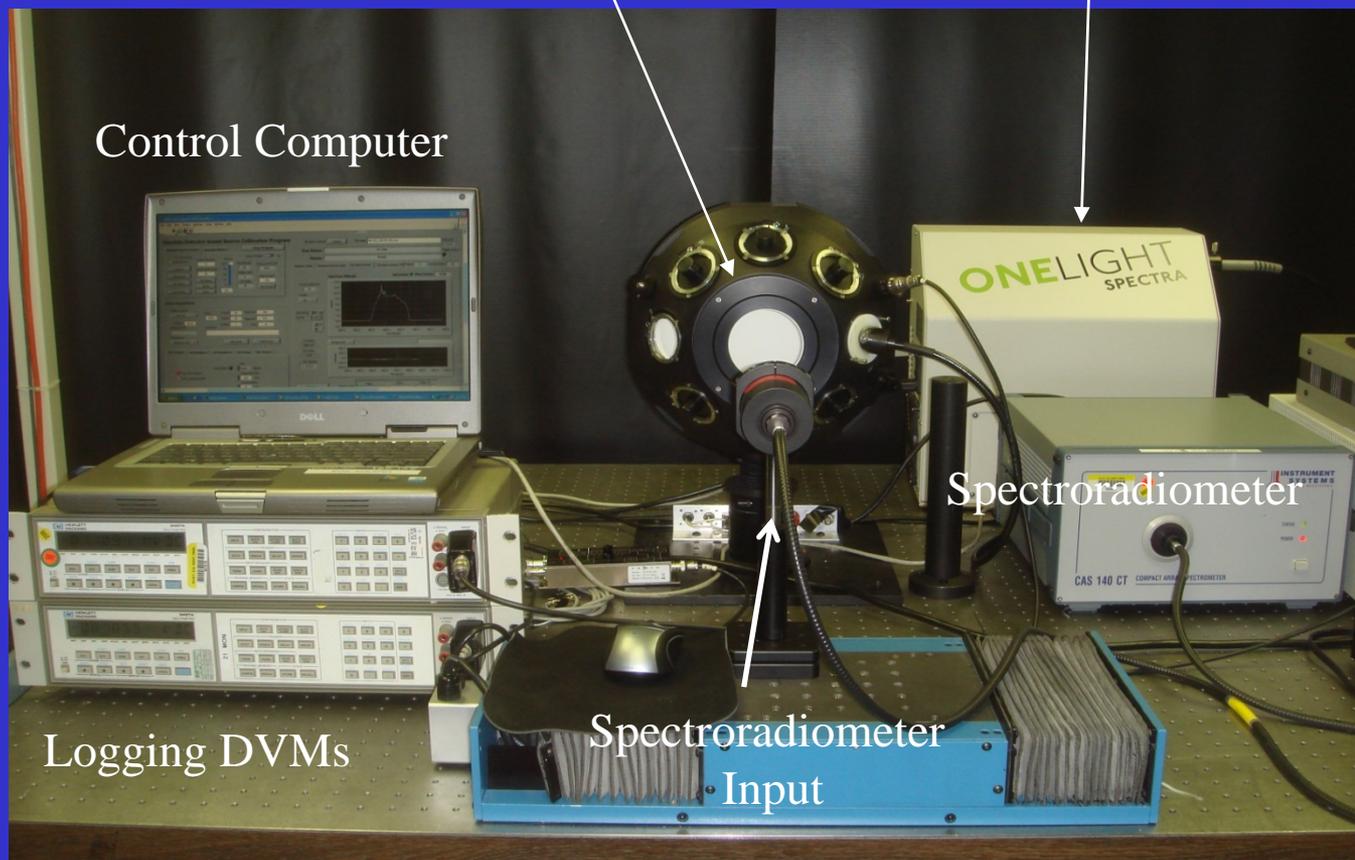
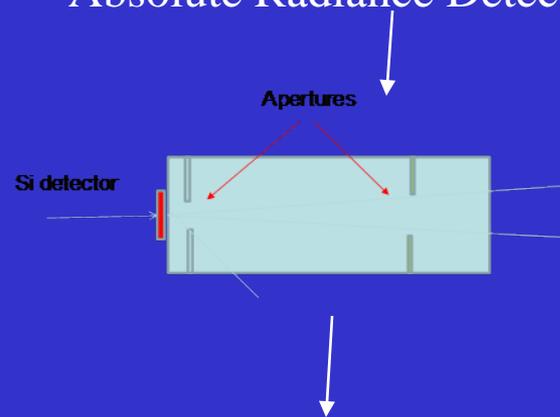
Integrating Sphere

Control Computer

Spectroradiometer

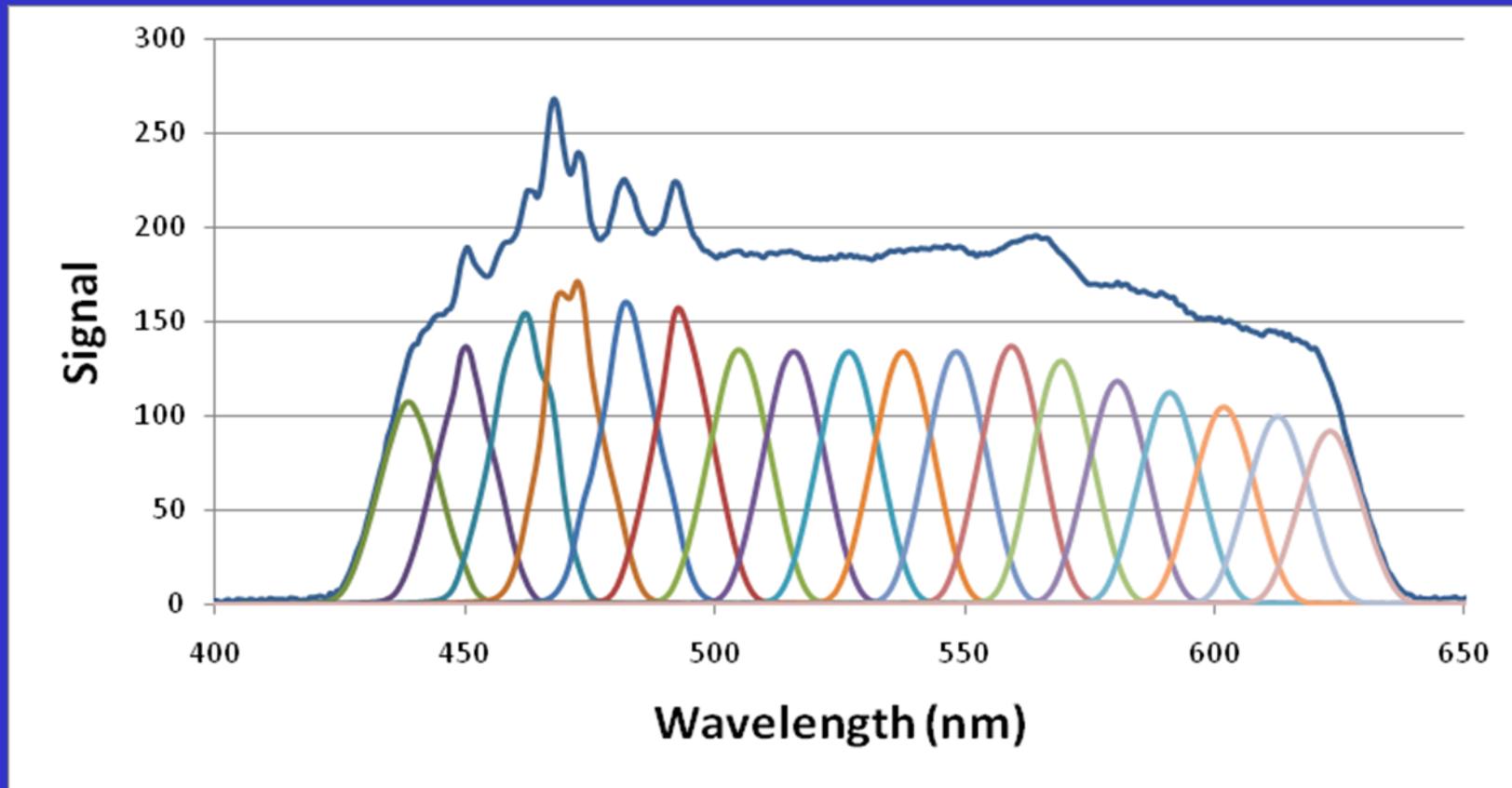
Logging DVMs

Spectroradiometer
Input



For more details, see Brown et al., *Proc. SPIE 7807*, 78070A (2010)

Broadband vs Narrowband Operation



For more details, see Brown et al., *Proc. SPIE 7807*, 78070A (2010)

Questions?