
Sampling and Analysis for GNSS Radio Occultation

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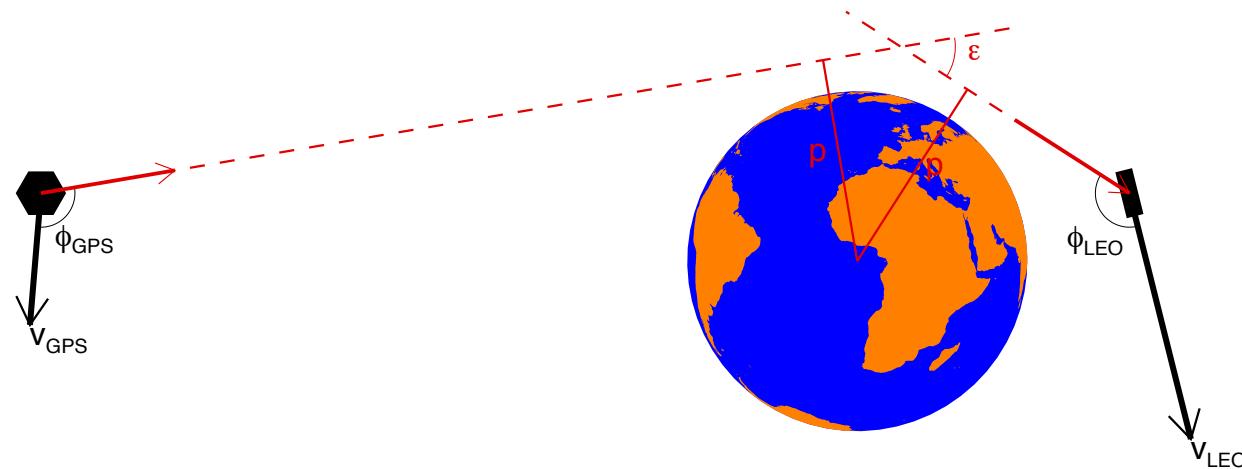
May 18, 2011

Outline

- Radio occultation (RO) on CLARREO
 - *SI traceability*
 - *Information content*
- Accuracy requirements
 - *Systematic vs. random*
 - *Sounding vs. climatological*
- Sampling error
 - *Systematic sampling error*
 - *Analysis: Bayesian interpolation*
 - *Early results*

GNSS Radio Occultation

$$-\frac{dL}{dt} = \lambda \Delta\nu = v_{\text{GPS}} \cos \phi_{\text{GPS}} + v_{\text{LEO}} \cos \phi_{\text{LEO}}$$



GNSS Radio Occultation (2)

- Orbit determination and clock correction, GPS and LEO: dL/dt
- Diffraction (and multipath) inversion: $\varepsilon(p)$

$$U(p) = \frac{1}{4\pi} \iint_S \left\{ U \frac{\partial}{\partial n} \left(\frac{e^{iks}}{s} \right) - \frac{e^{iks}}{s} \frac{\partial U}{\partial n} \right\} dS$$

- Inversion for refractivity: $N(r)$

$$\ln n(p) = \frac{1}{\pi} \int_p^{\infty} \frac{\varepsilon(p') dp'}{\sqrt{p'^2 - p^2}}$$

Kursinski, E.R., G.A. Hajj, J.T. Schofield, R.P. Linfield, and K.R. Hardy, 1997: Observing Earth's atmosphere with radio occultation measurements using the Global Positioning System. *J. Geophys. Res.*, **102**, 23429-23465.

Hajj, G.A., E.R. Kursinski, W.I. Bertiger, L.J. Romans, and S.S. Leroy, 2002: A technical description of atmospheric sounding by GPS occultation. *J. Atmos. Solar-Terr. Phys.*, **64**, 451-469.

Gorbunov, M.E., H.H. Benzon, A.S. Jensen, M.S. Lohmann, and A.S. Nielsen, 2004: Comparative analysis of radio occultation processing approaches based on Fourier integral operators. *Radio Sci.*, **39**, doi:10.1029/2003RS002916.

GNSS Radio Occultation (3)

- Refractivity

$$N = (n - 1) \times 10^6 = (77.6 \text{ K hPa}^{-1}) \frac{p}{T} + (363 \times 10^3 \text{ K}^2 \text{ hPa}^{-1}) \frac{p_w}{T^2}$$

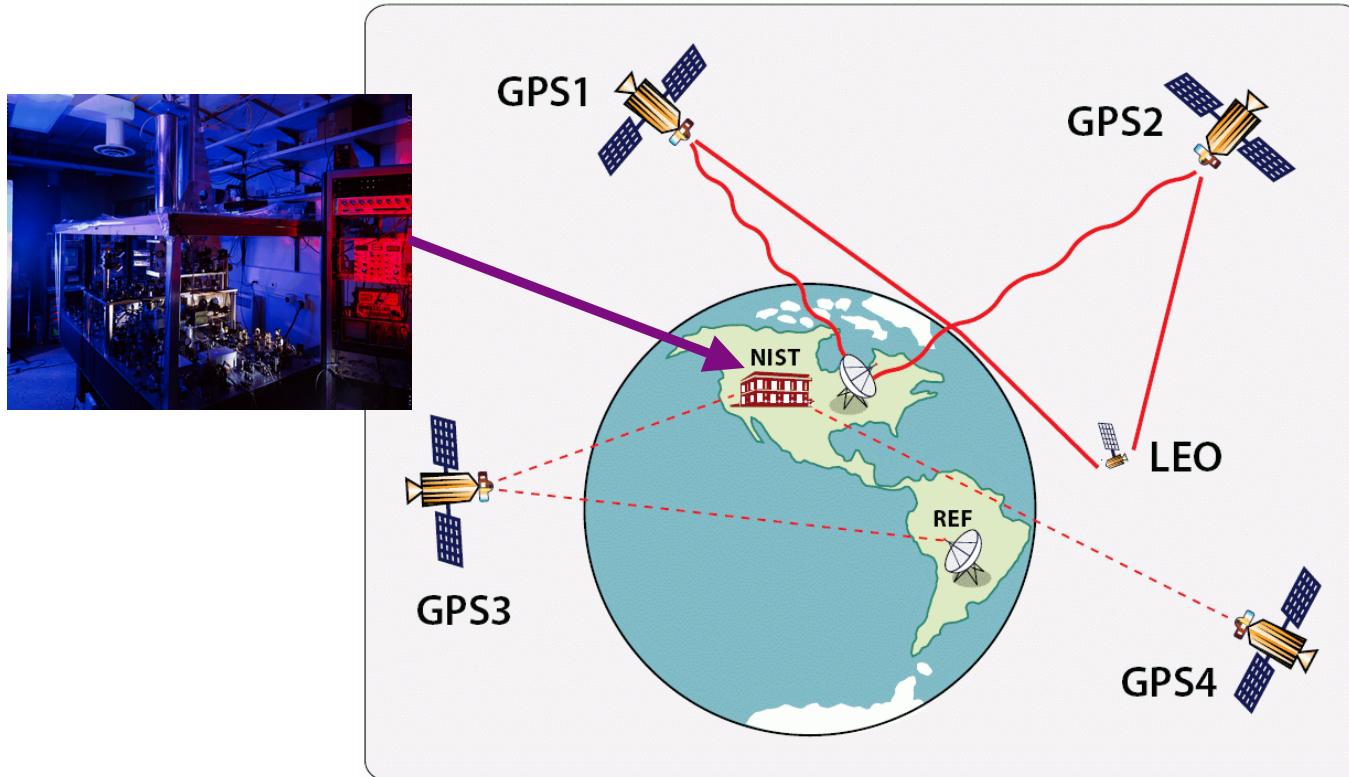
- “Dry” pressure

$$p_d(h) = (4.402 \times 10^{-4} \text{ hPa m}^{-1}) \int_h^{\infty} N dh \cong p(h) + (7521 \text{ K}) \int_0^{p(h)} \frac{q dp}{T}$$

- Geopotential height

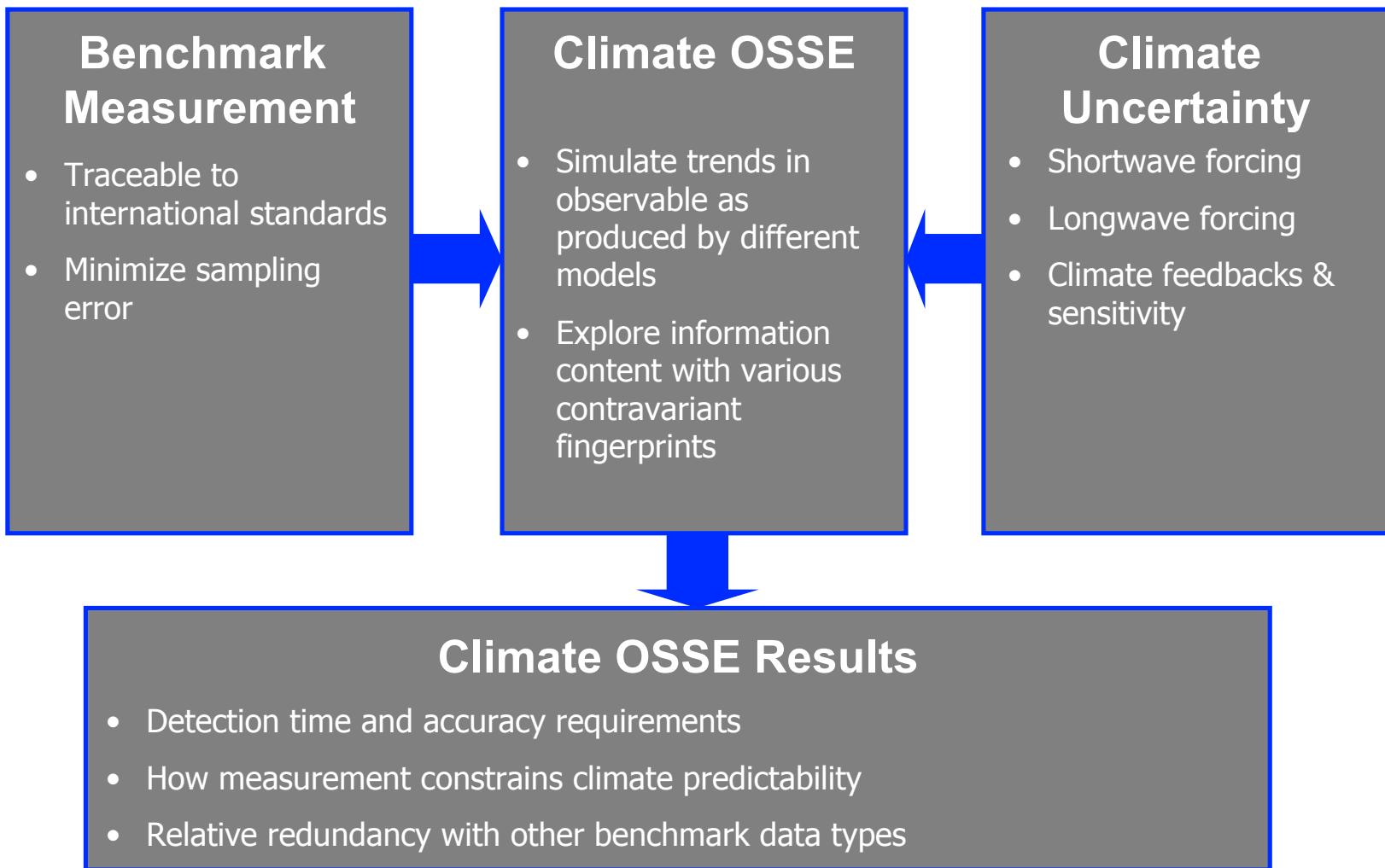
$$h = \left[(\Phi(\mathbf{r}) - \frac{1}{2} \Omega^2 r_s^2) - (\Phi - \frac{1}{2} \Omega^2 r_s^2)_{\text{msl}} \right] / g_0$$

GNSS Radio Occultation Calibration



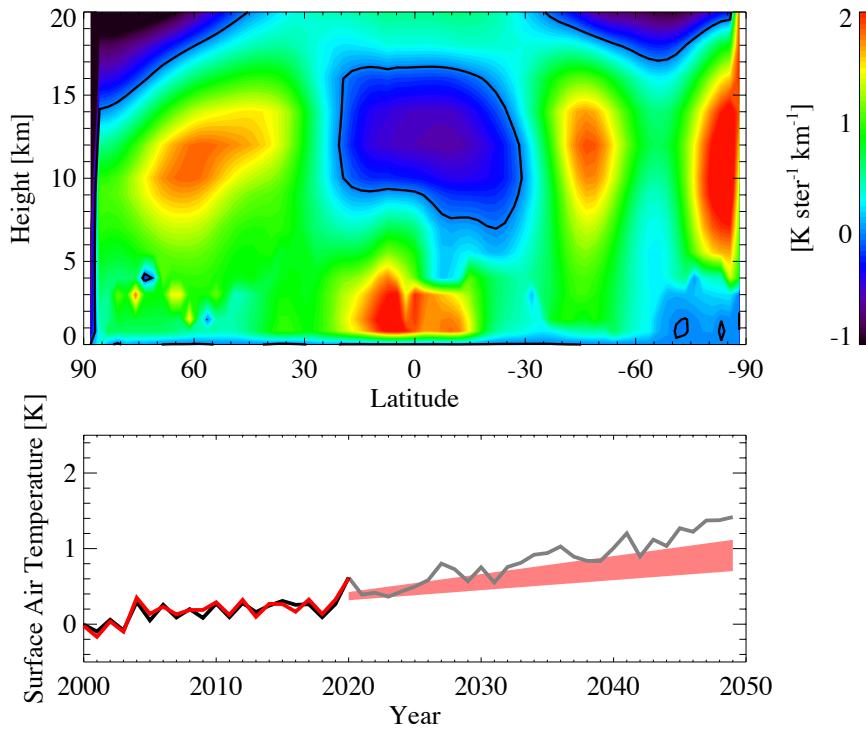
Hardy, K.R., G.A. Hajj, and E.R. Kursinski, 1994: Accuracies of atmospheric profiles obtained from GPS occultations. *Int. J. Sat. Comm.*, **12**, 463-473.

Climate OSSEs



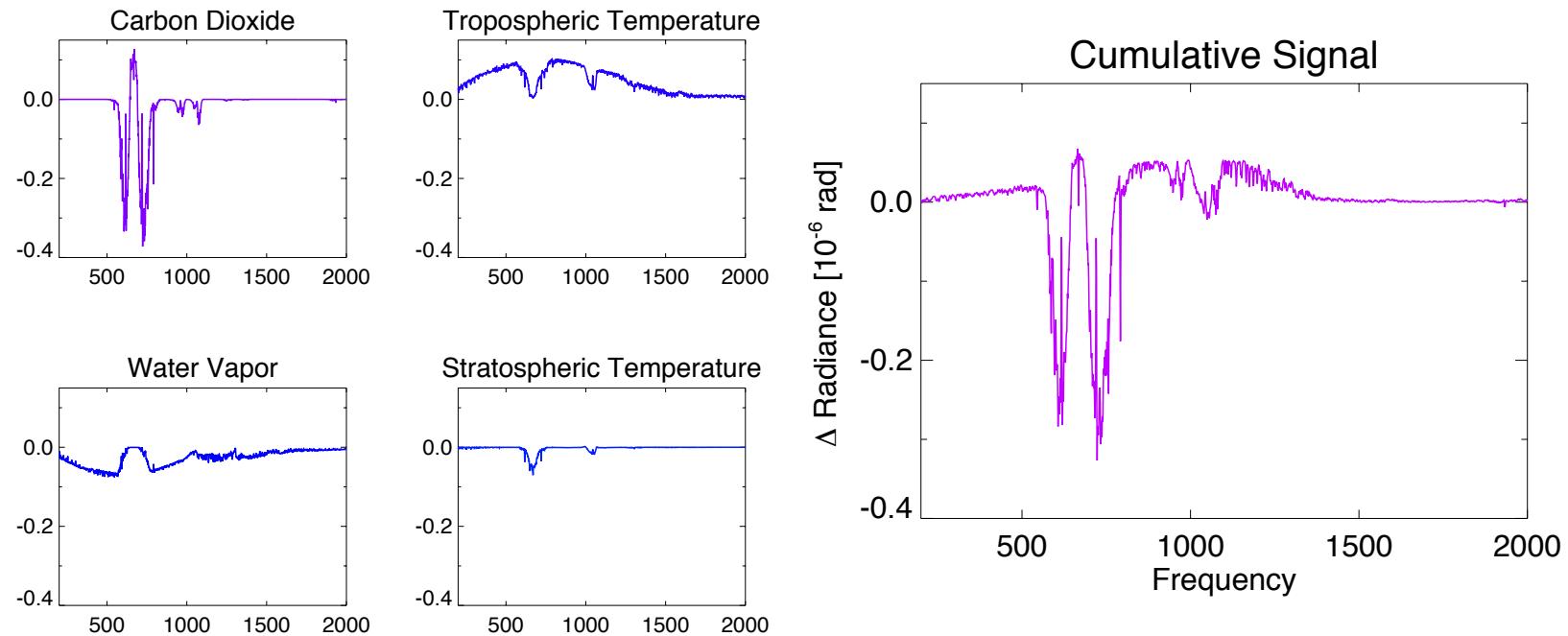
CLARREO GNSS Radio Occultation

- A substitute for surface air temperature monitoring.



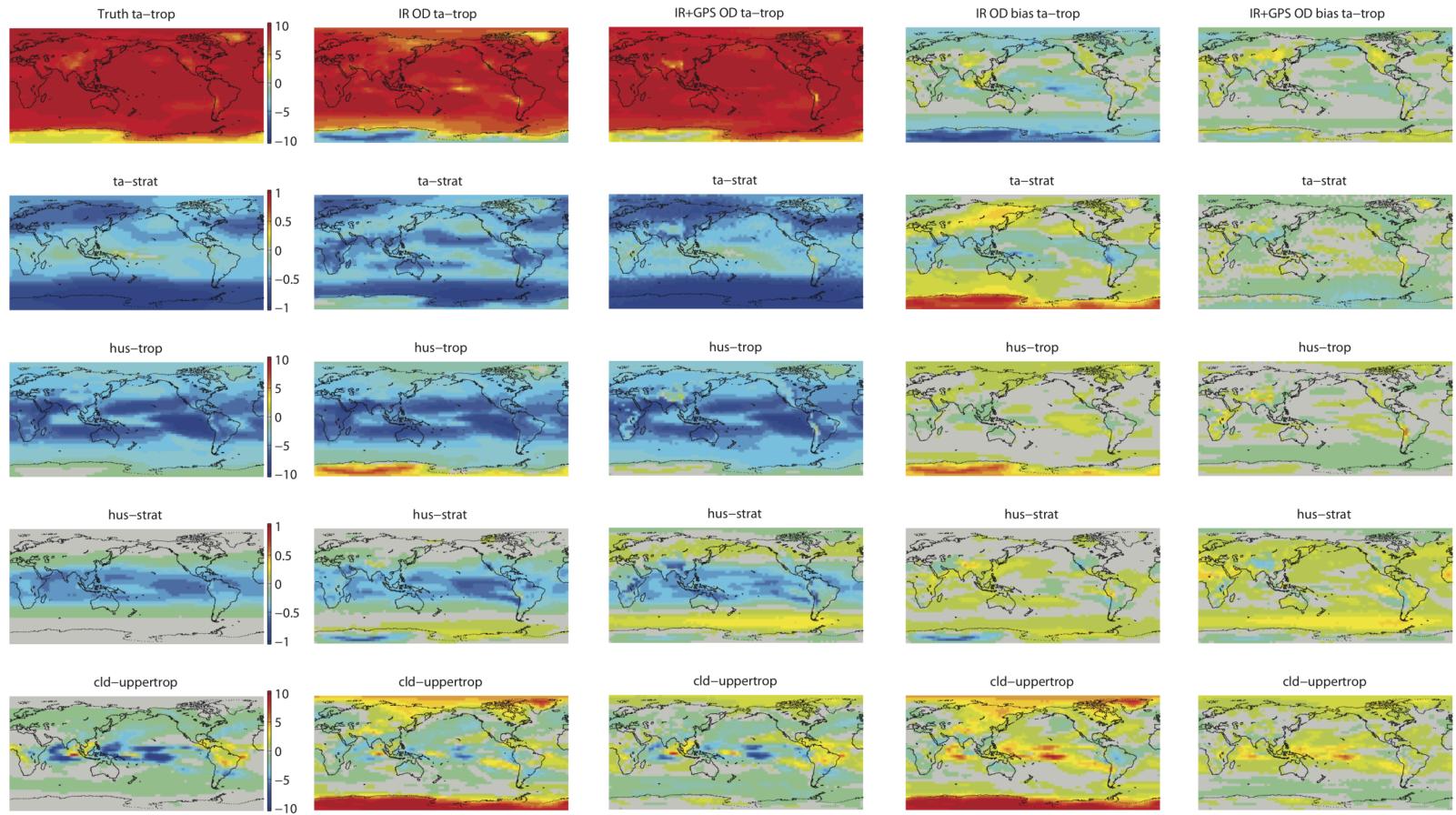
Leroy, Dykema, Gero, Anderson, 2009: Testing climate models using infrared spectra and GNSS radio occultation, in *New Horizons in Occultation Research*, Springer, 316pp.

Disentangling ambiguities in IR



Leroy, S.S., et al., 2008: Testing climate models using thermal infrared spectra. *J. Climate*, **21**, 1836-1848.

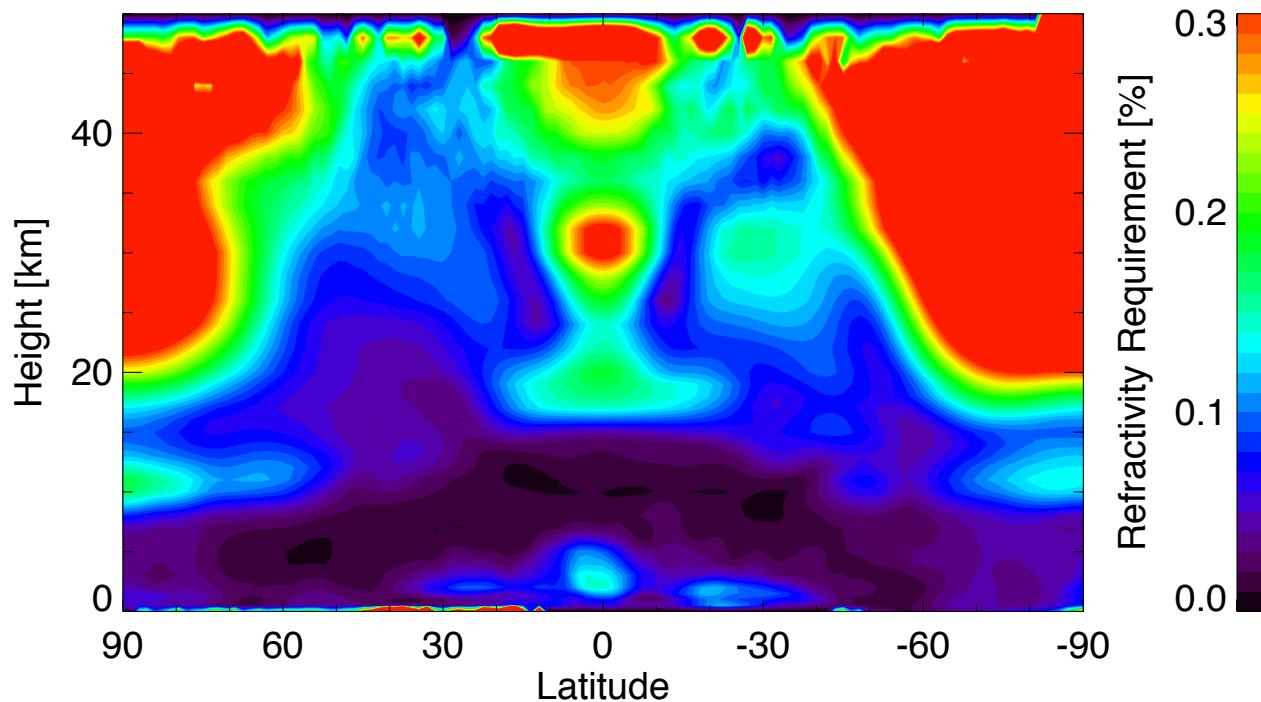
Disentangling ambiguities in IR (2)



Huang et al., 2011: Determining longwave forcing and feedback using infrared spectra and GNSS radio occultation.
J. Climate, In Press.

GNSS RO Accuracy Requirements

Accuracy requirements are set wherein total uncertainty does not increase time-to-detection more than 10% above what is dictated by natural variability.



Leroy, Anderson, Ohring, 2008: Climate signal detection times and constraints on climate benchmark accuracy requirements. *J. Climate*, **21**, 841-846.

Error Budget: GNSS RO

Refractivity requirement: 0.03% at 18 km, 5° bins.

Reviewed by the
RO community,
February 2, 2010.

CLARREO GNSS RO Error Budget				
Troposphere (5-20km)		Lower Troposphere (2-5km)		
Testing that requirements are met for this region to be done at 18km.		Testing that requirements are met for this region to be done at 3km.		
	Phase rate error (mm/s)	Refractivity error (%)	Phase rate error (mm/s)	Refractivity error (%)
Measurement Requirement A: Individual Sounding				
Systematic				
RO antenna phase center determination	0.021	0.0007%	0.021	0.0007%
Atmospheric multipath	0.000	0.000		
Ionospheric residual	0.300	0.0100%	0.060	0.0020%
LEO POD	0.054	0.0018%	0.000	0.0000%
Clock accuracy	0.030	0.0010%	0.000	0.0000%
Nonlinearity (retrieval?)			0.150	0.0050%
Attitude knowledge	0.024	0.0008%	0.024	0.0008%
Attitude rate knowledge	0.150	0.0050%	0.030	0.0010%
Local multi-path	0.240	0.0080%	0.060	0.0020%
<i>Total</i>	0.418	0.0139%	0.178	0.0059%
Random				
Instrument precision	0.240	0.0080%	0.060	0.0020%
Ionospheric scintillation	0.060	0.0020%	0.000	0.0000%
Gravity waves	0.090	0.0030%	0.000	0.0000%
Clock precision	0.099	0.0033%	0.000	0.0000%
<i>Total</i>	0.281	0.0094%	0.060	0.0020%
Measurement Requirement B: Climatological Averaging				
Systematic				
Diurnal cycle	n/a	0.0100%	n/a	0.0300%
<i>Total</i>	n/a	0.0100%	n/a	0.0300%
Random				
Sampling density	n/a	0.0220%	n/a	0.0800%
<i>Total</i>	n/a	0.0220%	n/a	0.0800%
<i>Total systematic error (target)</i>		0.0170%		0.0570%
Total systematic error		0.0172%		0.0306%
<i>Total random error (target)</i>		0.0250%		0.0820%
Total random error		0.0239%		0.0800%

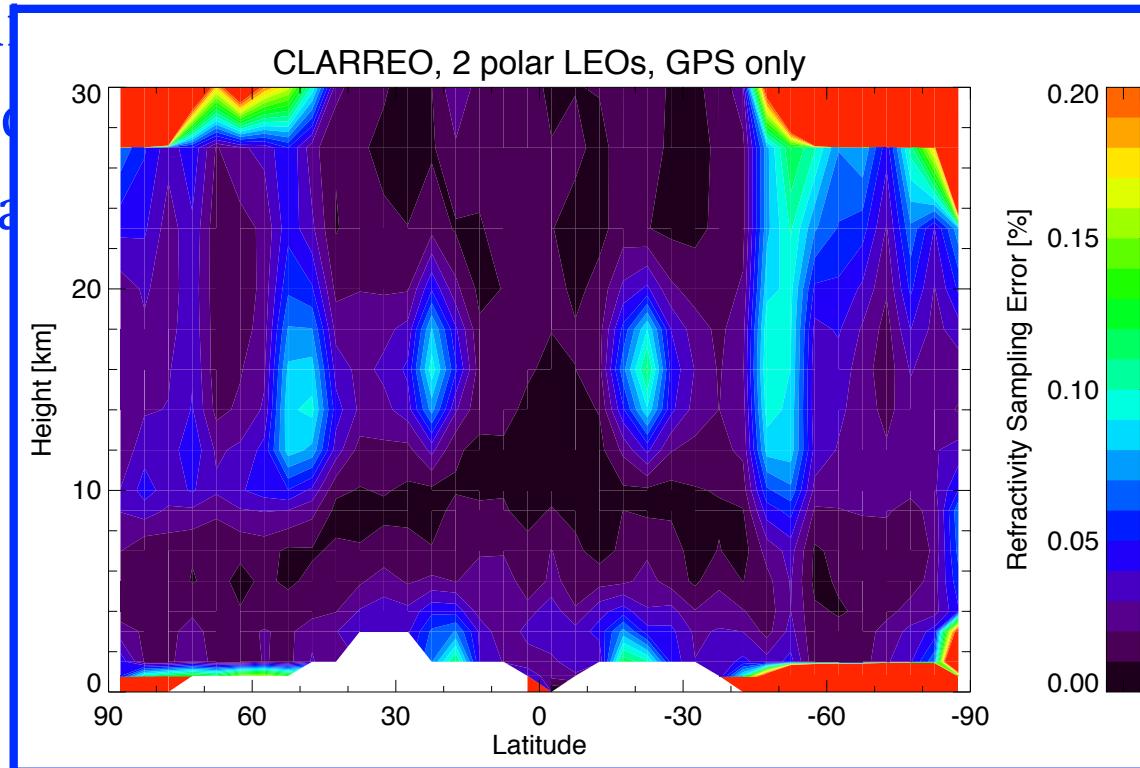
0.022%

Sampling Error

- Simulate a distribution of soundings.
- Interpolate reanalysis to time and location of soundings.
- Form climatology based on reanalysis “data”.
- Compare to reanalysis gridded “truth”.

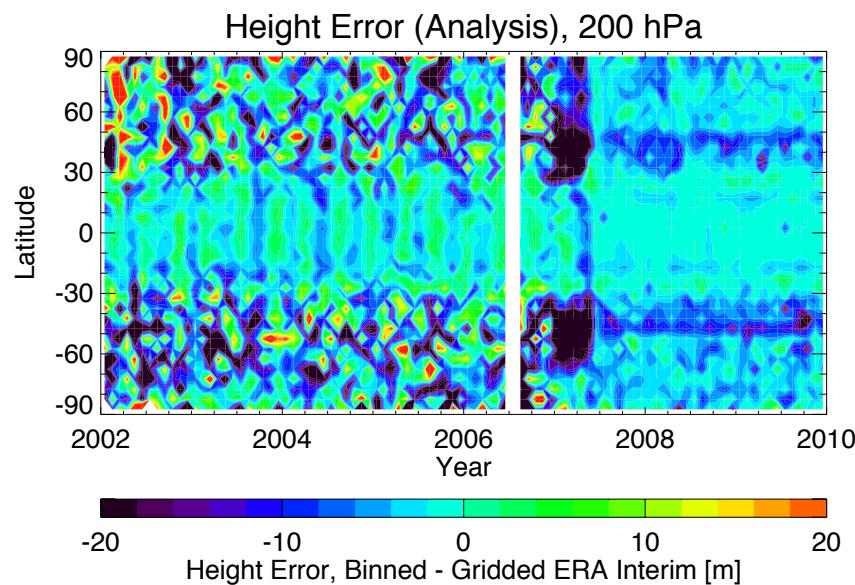
Sampling Error

- Simulate a distribution of soundings.
- Interpolate reanalysis to time and location of soundings.
- Form difference.
- Compare.



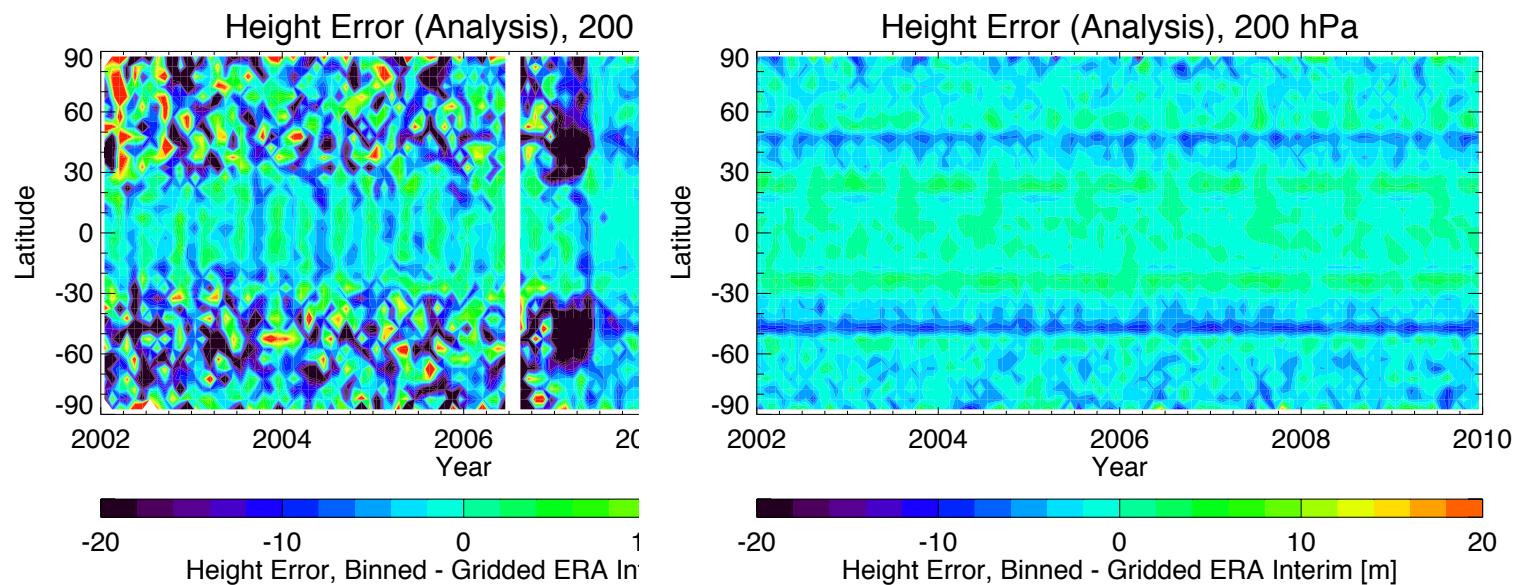
Sampling Error, CHAMP & COSMIC

Geopotential height of 200 hPa *dry* pressure surface



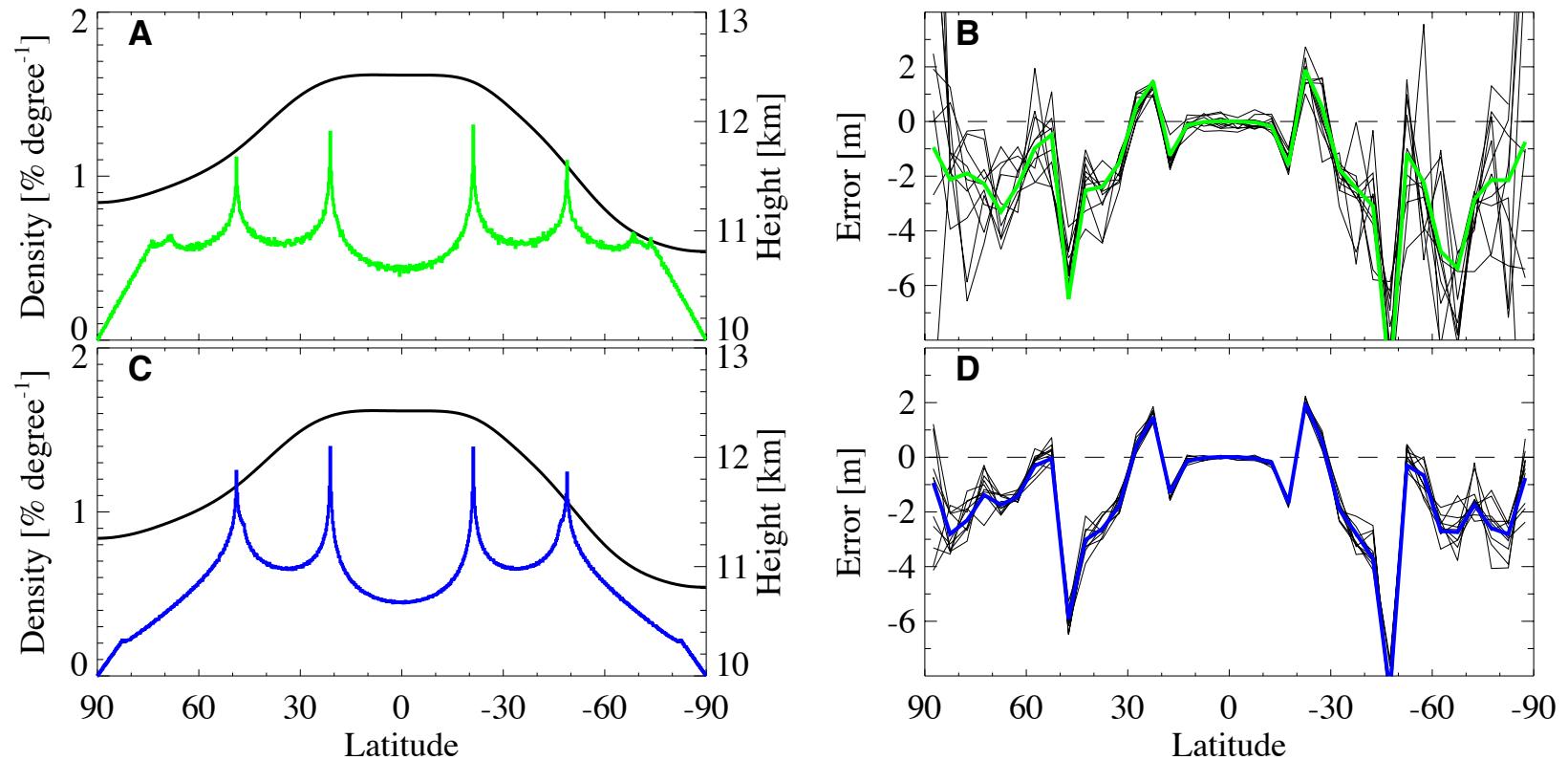
Sampling Error, CHAMP & COSMIC

Geopotential height of 200 hPa *dry* pressure surface

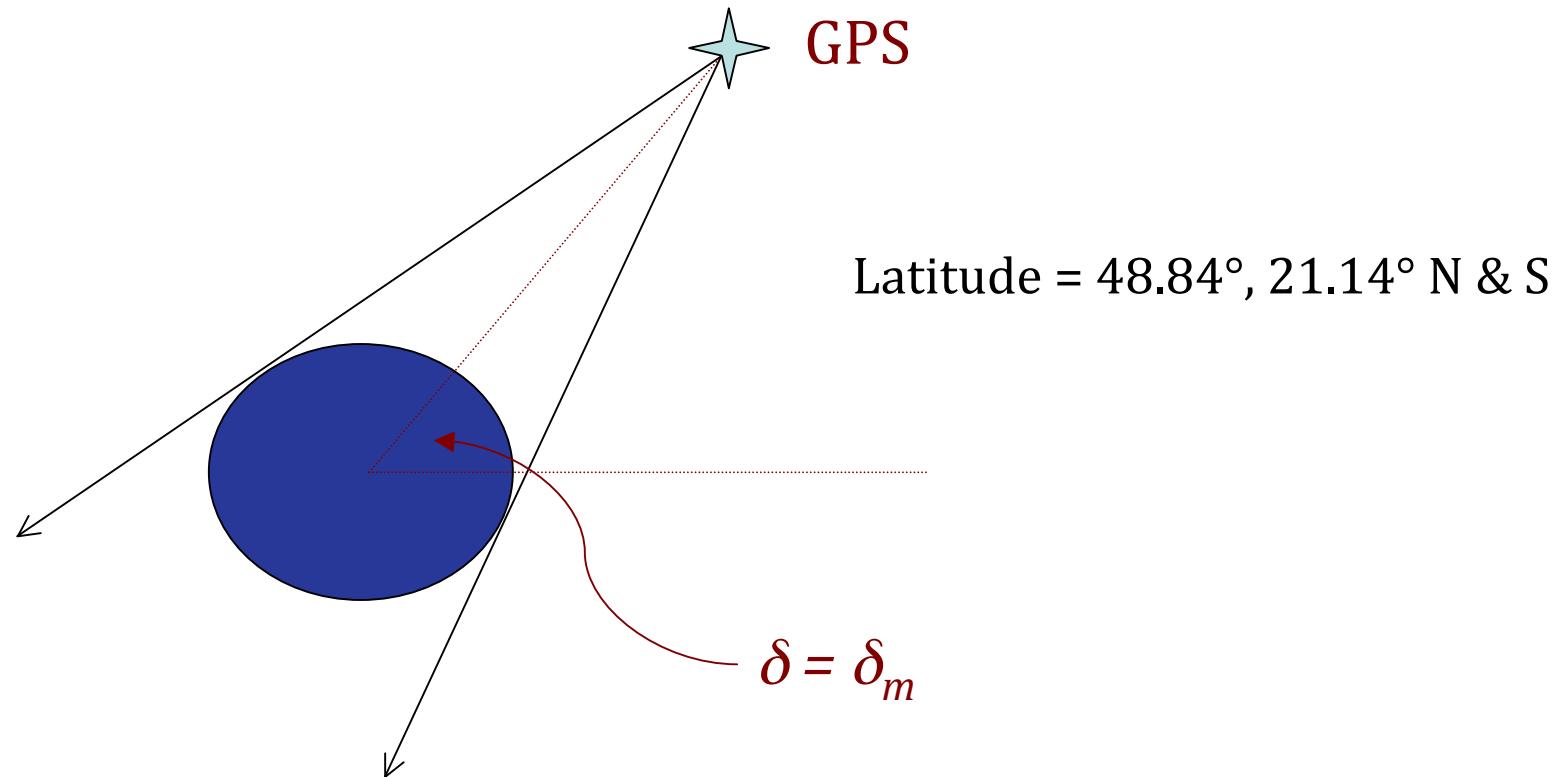


Sampling Error, CHAMP & COSMIC

Geopotential height of 200 hPa *dry* pressure surface



Systematic sampling error cause



Analysis: Bayesian Interpolation

Challenge: Fit data without under-fitting or overfitting.

Data under-samples unknown synoptic variability in space and time.

Solution: Use spherical harmonics as basis functions, apply hierarchical Bayesian inference to find best fit and most likely weighting of data fitting and penalty function.

Fit

Penalty

$$\chi^2 = \beta |\mathbf{t} - \phi \mathbf{w}|^2 + \alpha \mathbf{w}' \mathbf{C} \mathbf{w}$$

Analysis: Bayesian Interpolation (2)

$$\chi^2 = \beta |\mathbf{t} - \phi \mathbf{w}|^2 + \alpha \mathbf{w}' \mathbf{C} \mathbf{w}$$

$$\mathbf{B} = \phi \phi'$$

$$\mathbf{A} = \beta \mathbf{B} + \alpha \mathbf{C}$$

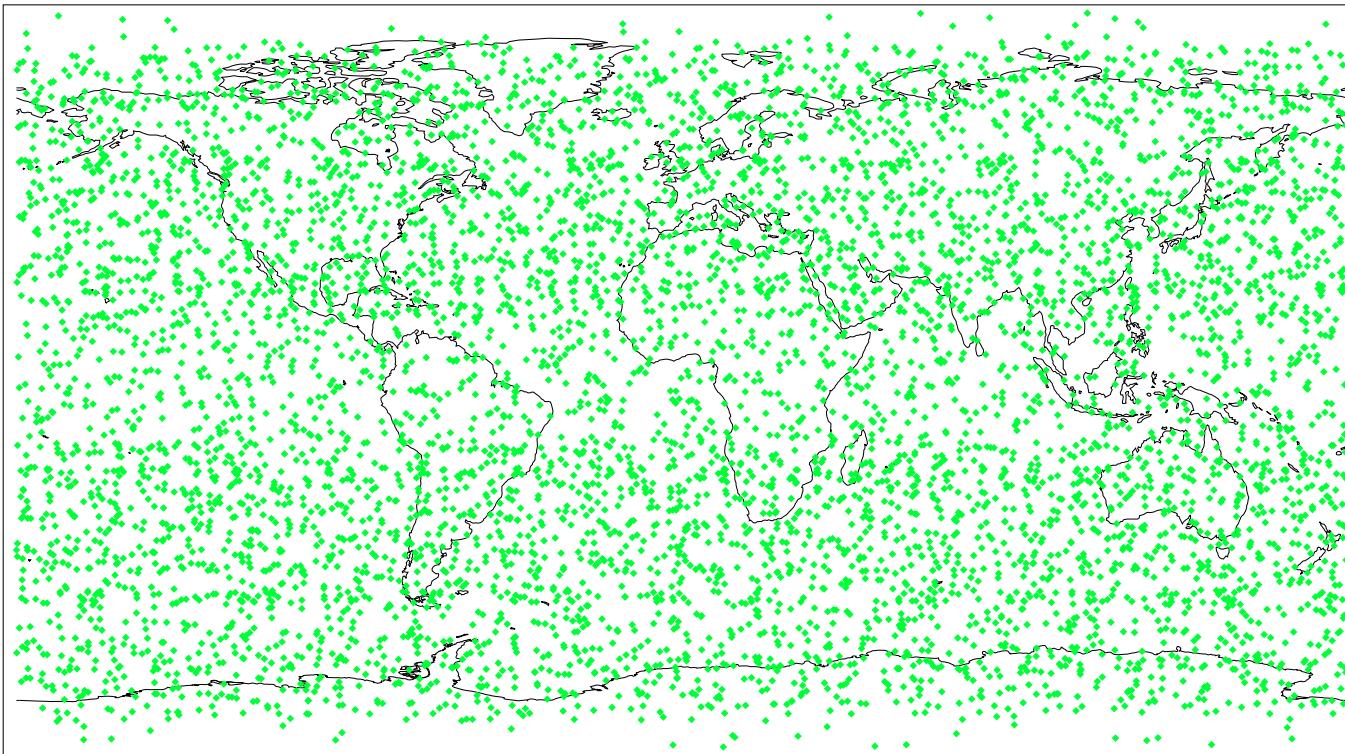
First inference:

$$\mathbf{w} = \mathbf{A}^{-1} \beta \phi' \mathbf{t}$$

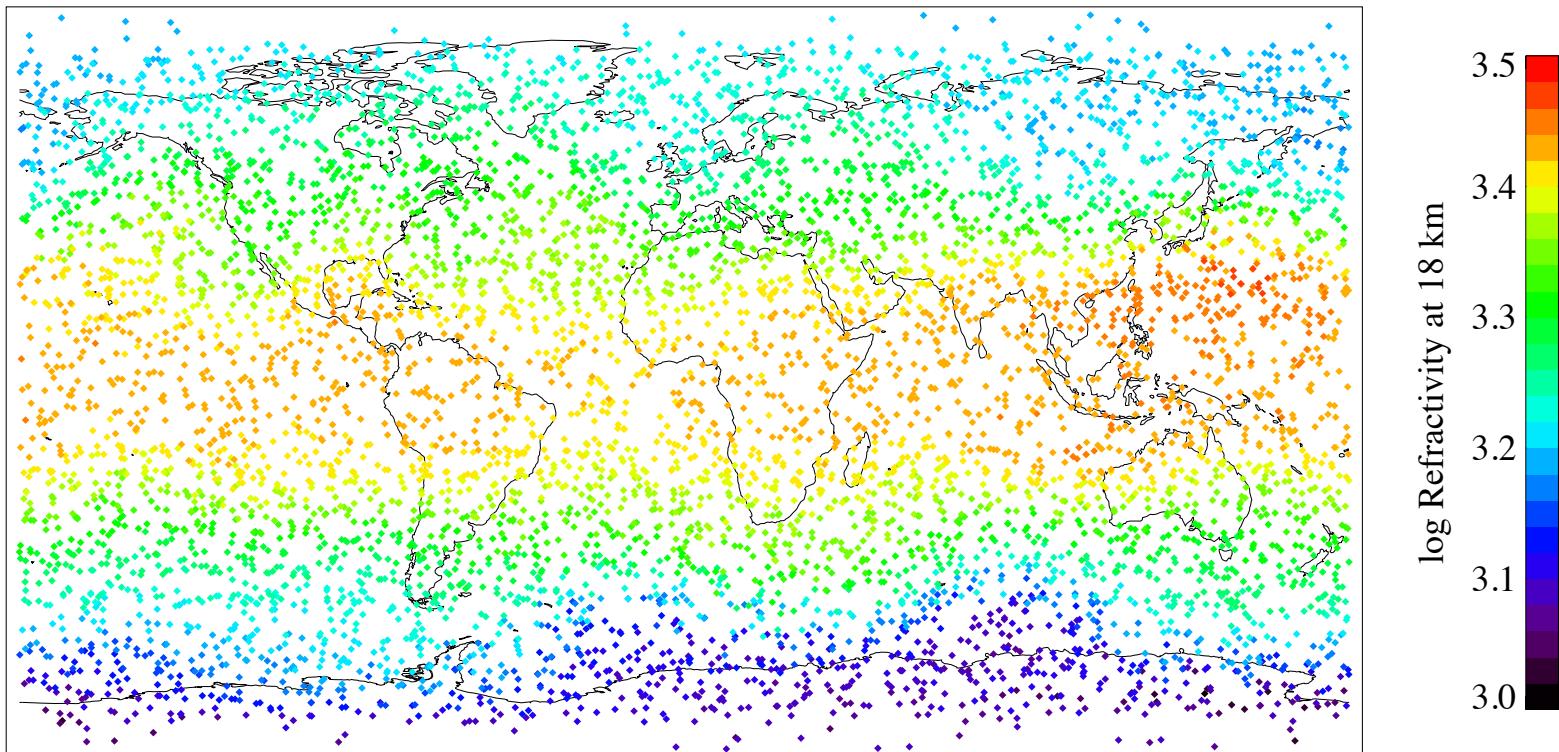
Second inference:

$$\gamma = k - \alpha \text{Trace } \mathbf{A}^{-1} \mathbf{C} = N - \beta |\mathbf{t} - \phi \mathbf{w}|^2$$

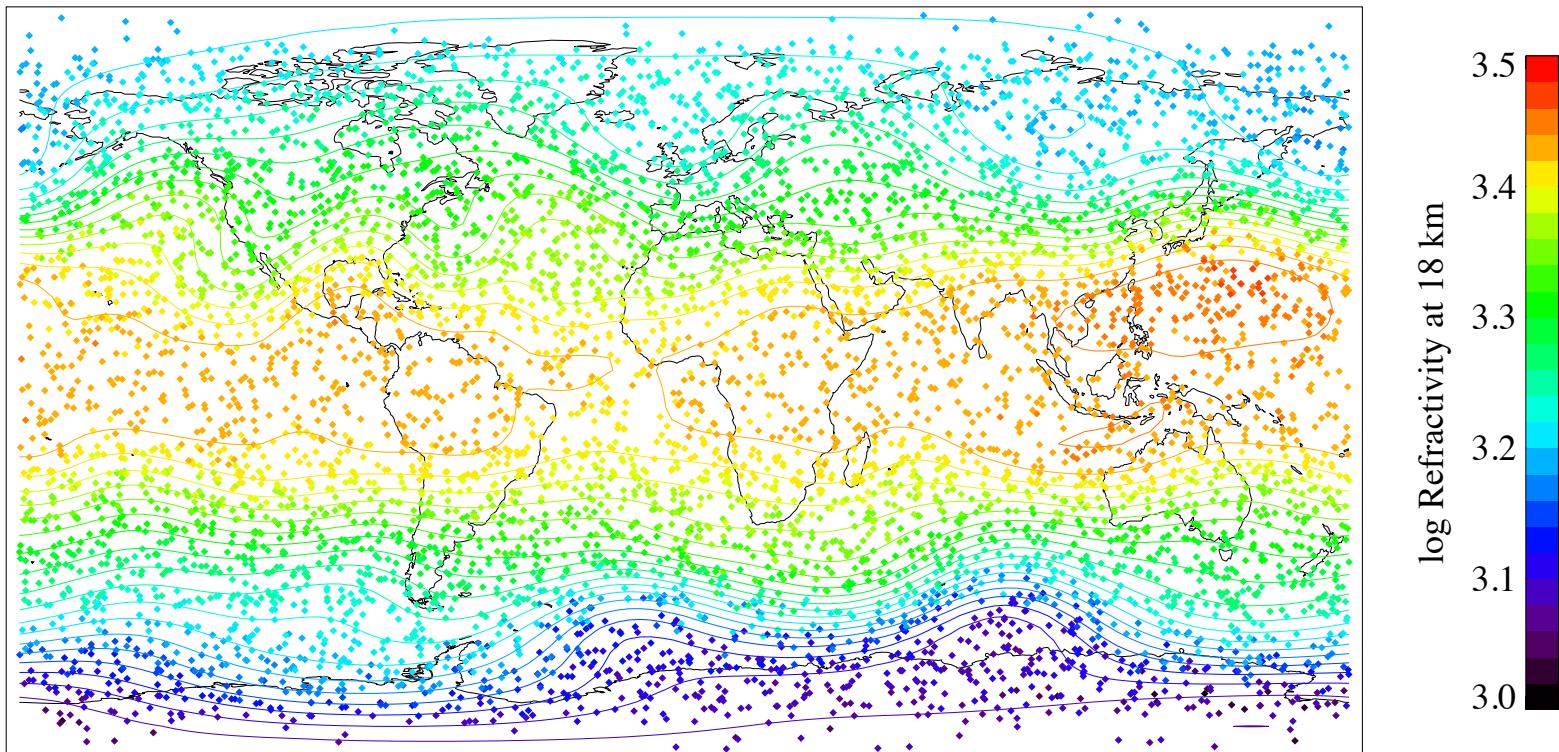
Analysis: Bayesian Interpolation (3)



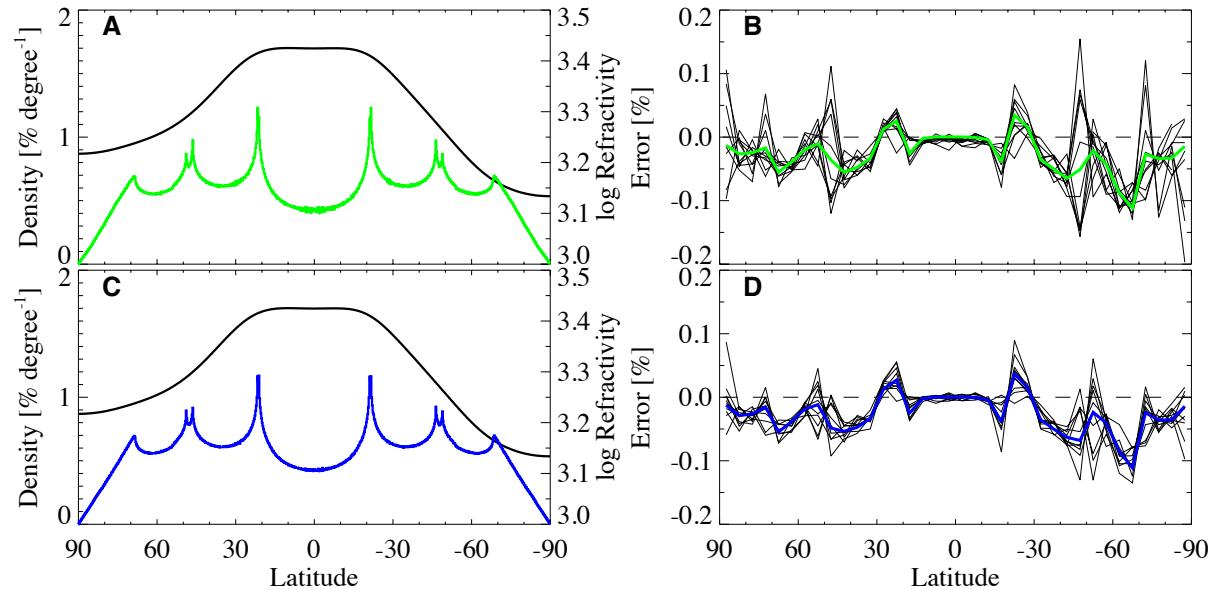
Analysis: Bayesian Interpolation (3)



Analysis: Bayesian Interpolation (3)

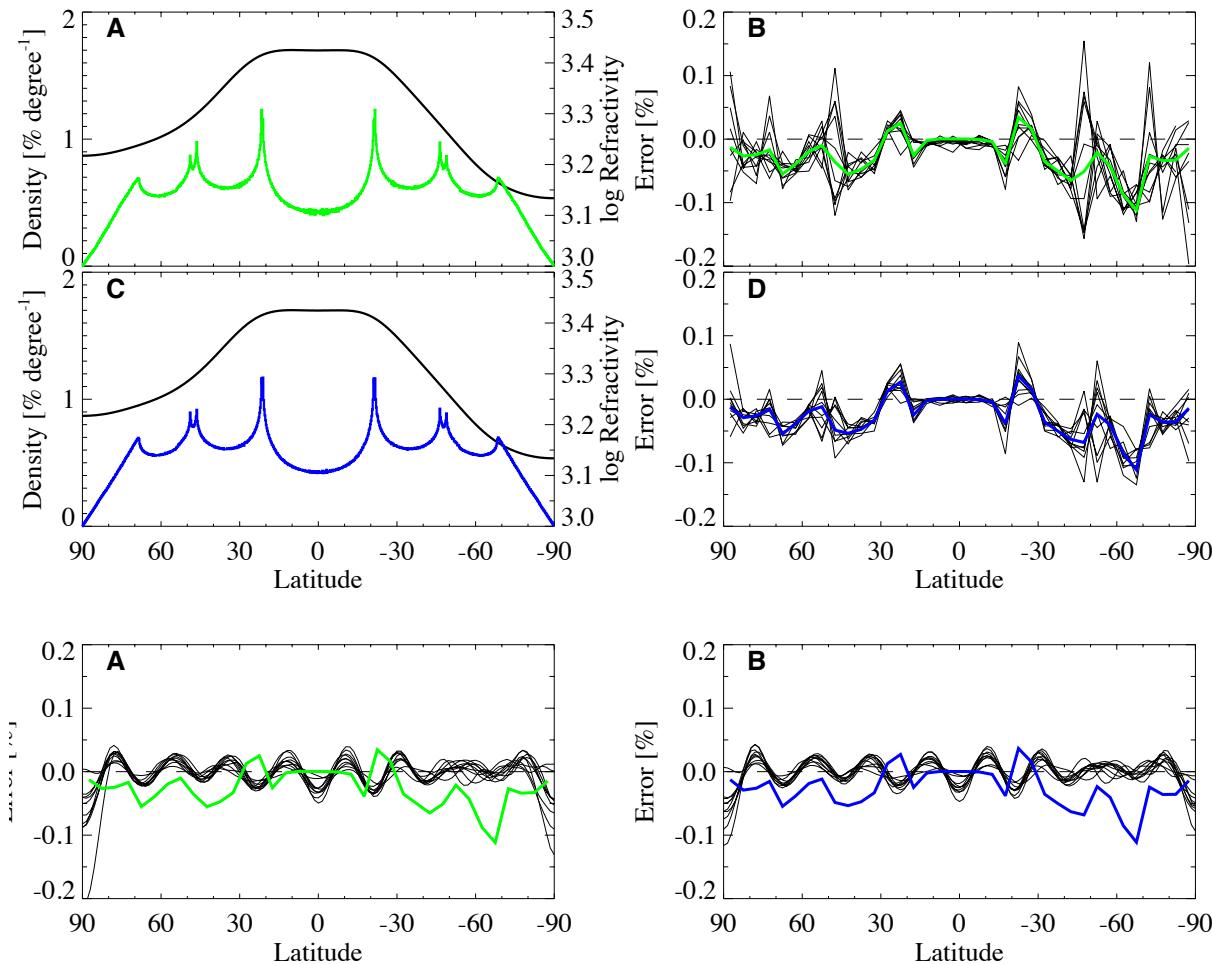


Systematic Sampling Error



Binning and
averaging

Systematic Sampling Error



Binning and averaging

Bayesian interpolation

Summary

- A systematic sampling error arises due to weak singularities in RO coverage. The singularities are at 48° and 21° north and south latitude.
- Bayesian interpolation greatly reduces systematic error and random error. Residual systematic error due to spherical harmonic truncation.
- CLARREO sampling requirements will clearly be effected, most likely downward. Yet to determine temporal sampling errors and consequently the number of required CLARREO orbits.

This work was funded by a grant from the NASA Jet Propulsion Laboratory's Director's Research and Discretionary Fund and by the CLARREO Project.

Extra slides

GNSS RO Missions

- **GPS/MET**, UCAR, 1995-1997. Soundings in four “prime time” periods.
- **CHAMP**, GFZ Potsdam, 2001-2009. ~200 soundings daily.
- **MetOp-A**, EUMETSAT, 2006-present. ~500 soundings daily; intermittent availability.
- **COSMIC**, UCAR, 2006-present. ~2800 soundings daily; degrading because of age.
- **TerraSAR-X**, DLR, 2008-present. Presently unavailable.

More general: Bayes (or just marginalization)

$$P(\mathbf{x}, \mathbf{y}) = P(\mathbf{x} | \mathbf{y}) P(\mathbf{y}) = P(\mathbf{y} | \mathbf{x}) P(\mathbf{x})$$

\mathbf{x} : observable variables

\mathbf{y} : prediction variables

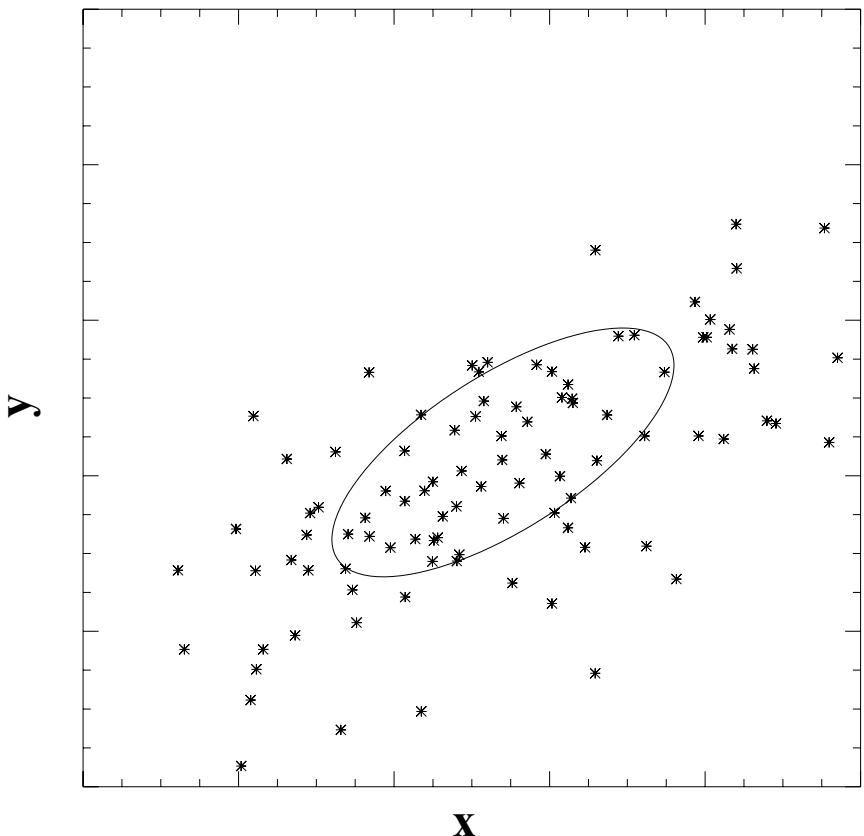
Form joint PDF $P(\mathbf{x}, \mathbf{y})$ using an ensemble of climate models. For each model, need (1) observation kernel to simulate data \mathbf{x} from hindcast run, and (2) emissions scenario run to generate prediction variables \mathbf{y} .

Natural variability in \mathbf{x} and \mathbf{y} and uncertain physics will both be accounted for.

With data \mathbf{d} , set $\mathbf{x} = \mathbf{d}$ and $P(\mathbf{y}|\mathbf{x}=\mathbf{d})$ is the projection PDF with data incorporated. $P(\mathbf{x})$ is a normalization constant that guarantees a unit integral of $P(\mathbf{y}|\mathbf{x})$ over \mathbf{y} .

Simple Solution

Assume Gaussian statistics



Covariance for (x,y) :

$$\Sigma = \begin{bmatrix} \sigma_x^2 & \rho\sigma_x\sigma_y \\ \rho\sigma_x\sigma_y & \sigma_y^2 \end{bmatrix}$$

Projection y :

$$\bar{y}_{\text{posterior}} = \bar{y}_{\text{prediction}} + \rho \frac{\sigma_y}{\sigma_x} d$$

Projection uncertainty:

$$\sigma_{y,\text{posterior}}^2 = \sigma_y^2 (1 - \rho^2)$$

Multivariate Solution

$$\Sigma = \begin{bmatrix} \Sigma_{xx} & \Sigma_{xy} \\ \Sigma_{yx} & \Sigma_{yy} \end{bmatrix}$$

Correlation of prediction and data

$$\mathbf{y}_{\text{posterior}} = \bar{\mathbf{y}}_{\text{prior}} + \Sigma_{yx} \Sigma_{xx}^{-1} (\mathbf{d} - \bar{\mathbf{x}}_{\text{prior}})$$
$$\Sigma_{y|d} = \Sigma_{yy} - \underbrace{\Sigma_{yx} \Sigma_{xx}^{-1} \Sigma_{xy}}$$

Information provided by data

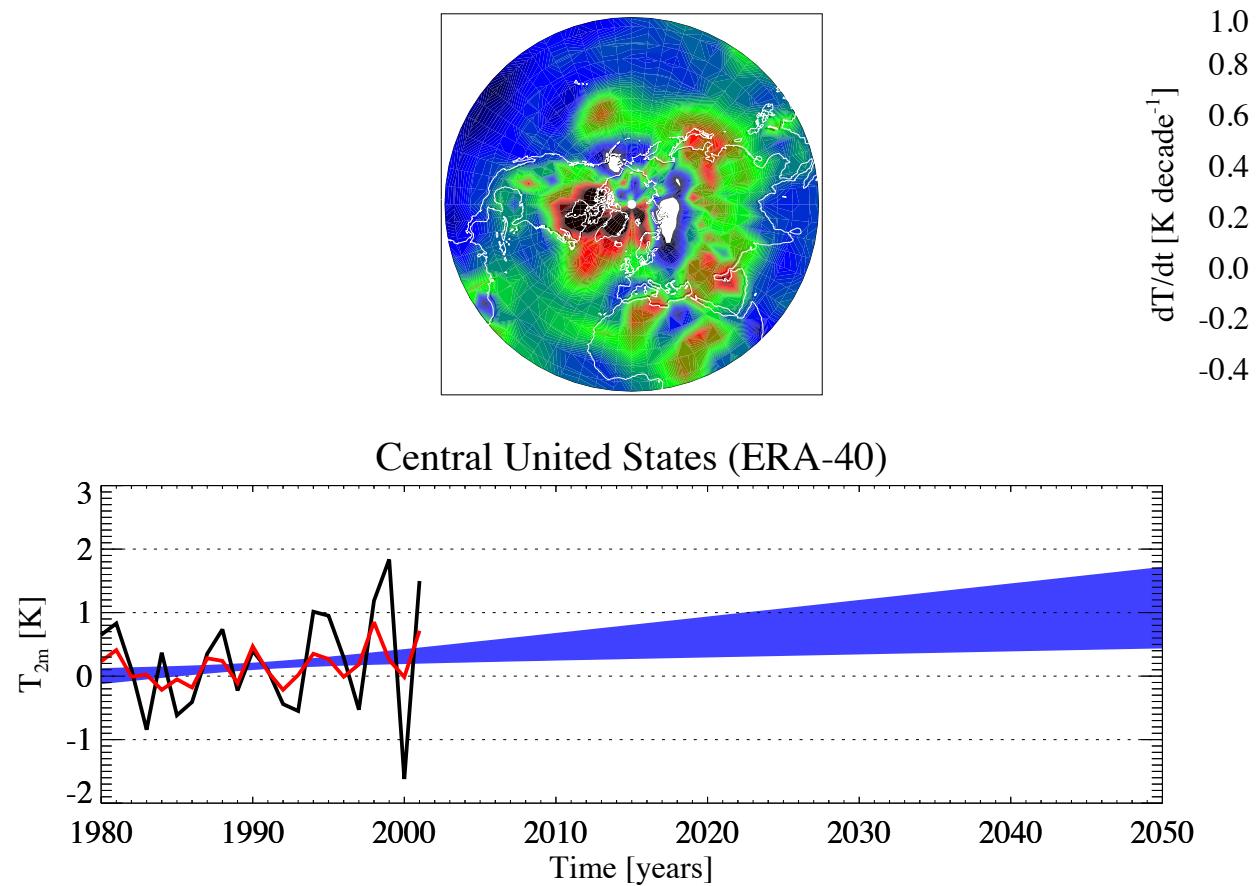
Ensemble uncertainty *Optimization*

- Continuum of models is assumed, solving the problem of reality ensemble averaging (REA)
- Higher dimensionality in x (observation) reduces posterior uncertainty but requires greater number of models in ensemble.
- Information is provided by correlation. *Which data types are the most important for improving climate prediction?*

Information in data types

Infer 50-yr temperature trend from observed 20-yr trends:

Application of Reanalyses: Central U.S.



Application of Reanalyses: Sahel

